

VLBI2010 Project for Geodesy and Astrometry

Arthur Niell

The VLBI2010 Committee

The Broadband Development Team

VLBI2010 Committee

Bill Petrachenko¹ (chair), Arthur Niell², Dirk Behrend³, Brian Corey²,
Johannes Böhm⁴, Patrick Charlot⁵, Arnaud Collioud⁵, John Gipson³,
Rüdiger Haas⁶, Thomas Hobiger⁷, Yasuhiro Koyama⁷, Dan MacMillan³,
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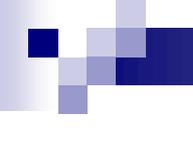
⁵*Bordeaux Observatory, France*

⁶*Onsala Space Observatory, Chalmers University of Technology, Sweden*

⁷*Kashima Space Research Center, NICT, Japan*

⁸*Pulkovo Observatory, Russia*

⁹*Radio Astronomy Institute, Italian National Astrophysical Institute, Italy*



Broadband Delay Team

Bruce Whittier¹, Mike Titus¹, Jason SooHoo¹,
Dan Smythe¹, Alan Rogers¹, Jay Redmond²,
Mike Poirier¹, Arthur Niell¹, Chuck Kodak³,
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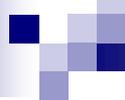
Special thanks to Sandy Weinreb and Hamdi Mani

VLBI2010 Recommendations

- 1-mm position accuracy on global scales
- Continuous measurements for time series of station positions and Earth orientation parameters
- Turnaround time to initial geodetic results of less than 24 hours

Principal VLBI2010 Actions

- Numerical simulations
- Proof-of-Concept Development
- Progress report (through end of 2008)
 - Completed (last week!)
 - Available soon on IVS web site
<http://ivscg.gsfc.nasa.gov>



Numerical Simulation Innovations

- More accurate atmosphere model that includes turbulence properties
- New approach to scheduling observations within a day
- Evaluation of effect of source structure

Recommendations from Simulations

- Observe as many sources as possible per hour.
- Antenna diameter 12 m or greater
- Antenna slew rates
 - if only one antenna on site, higher than 10 /sec in az and 3 /sec in el
 - if possibly two antennas on site, ~5 /sec in az and 1.5 /sec in el
- Initially the frequency range should be 2.2 GHz to ~14 GHz.

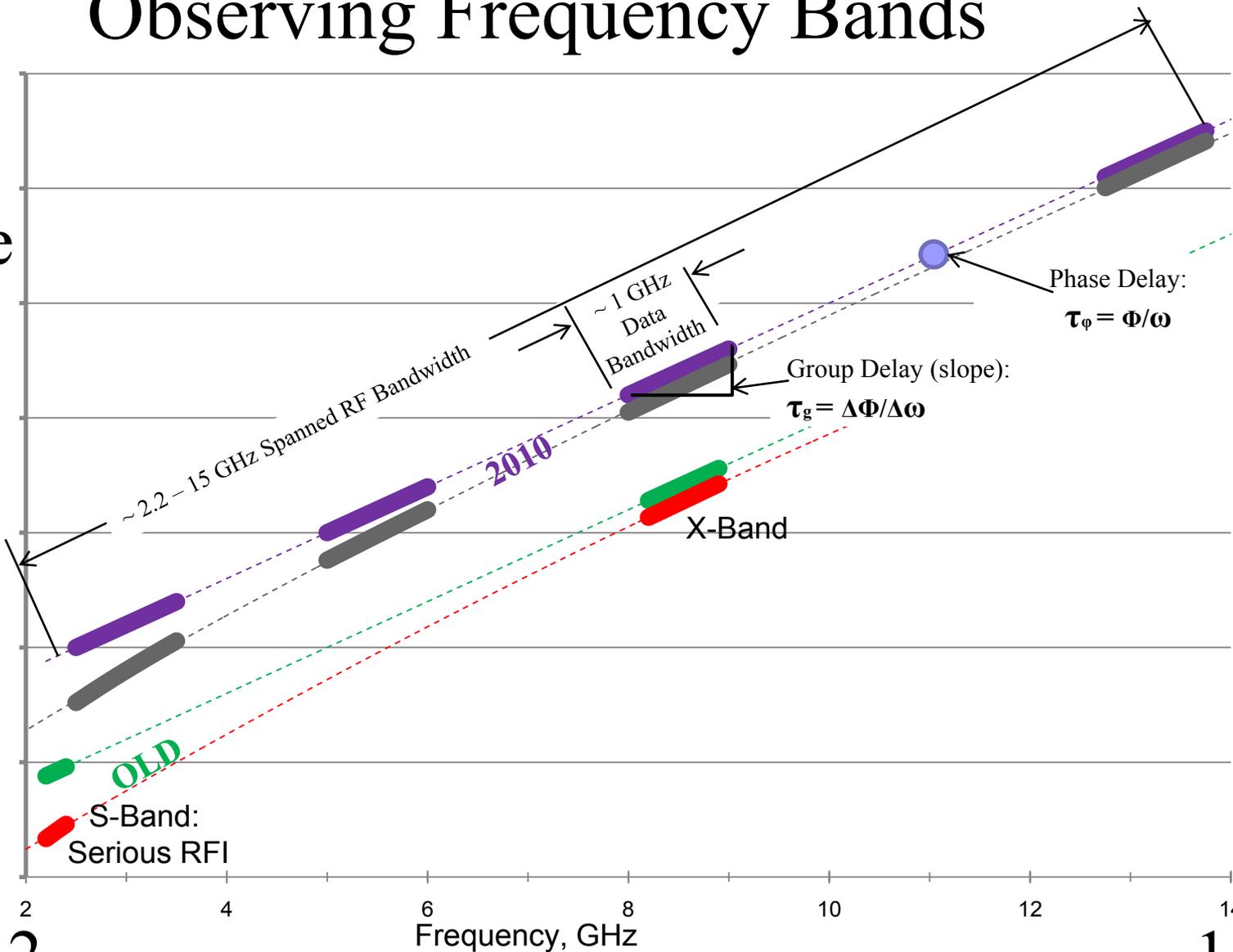
Proof-of-Concept Development

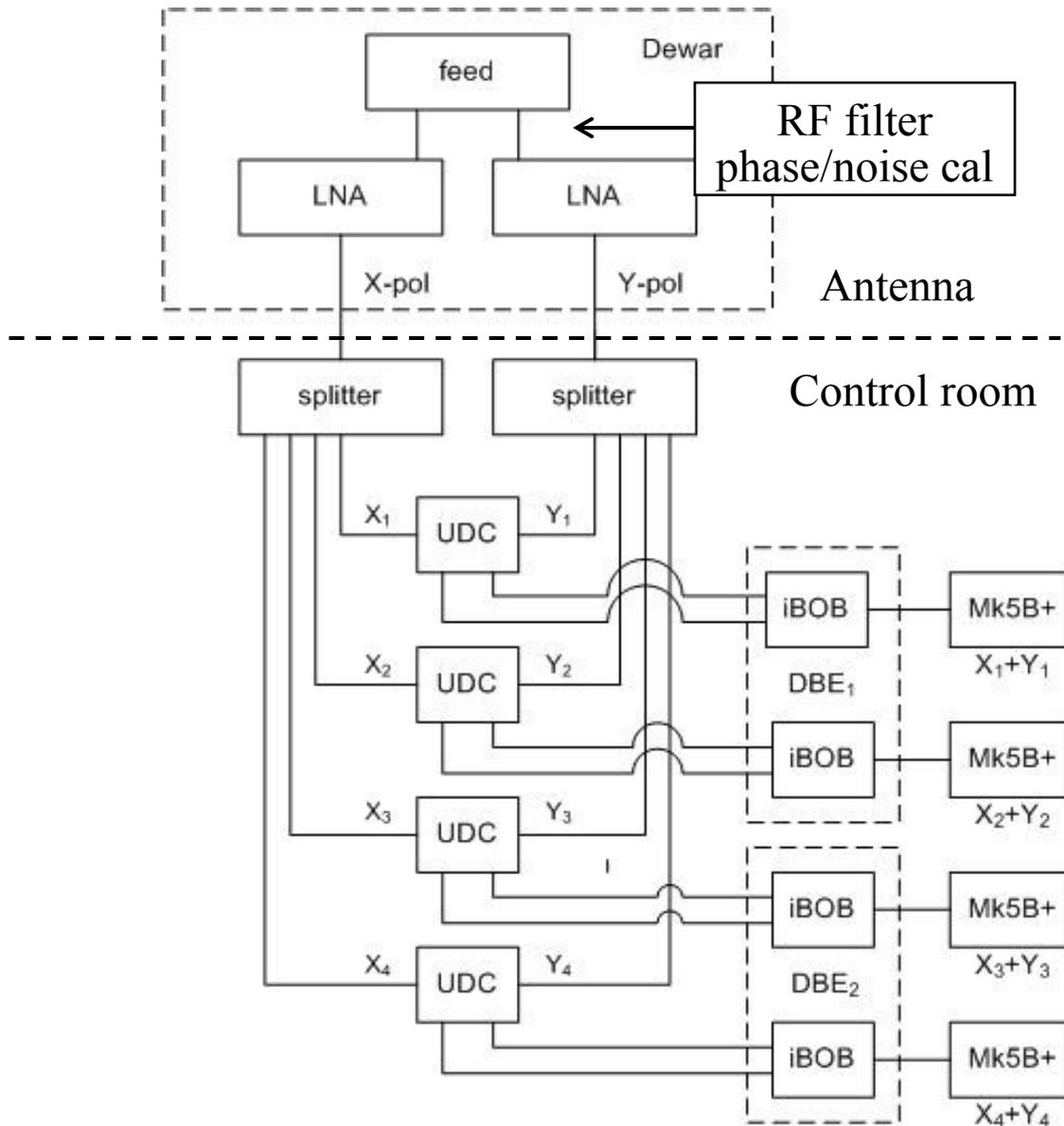
- Develop the instrumentation to implement the broadband delay system as recommended by the V2C in the IVS WG3 Report
- Mount the equipment on two existing antennas
 - Westford 18m antenna, Massachusetts
 - MV-3 5m antenna, Washington, D.C.
- Make observations to demonstrate that broadband delay can be effectively used

Observing Frequency Bands

Phase

PHASE(Arbitrary Units)





Feed and LNAs cooled to ~20K

Both senses of linear polarization used

Odd channels from each pol'n for one band output to each Mk5B+.

2 gigabits/sec recorded on each Mk5B+.

Total data rate: 8 gbps

MV-3 5M Antenna @ GGAO

Cryo LNAs & Feed
in Dewar



8 gigabit/sec
LOs and
back end



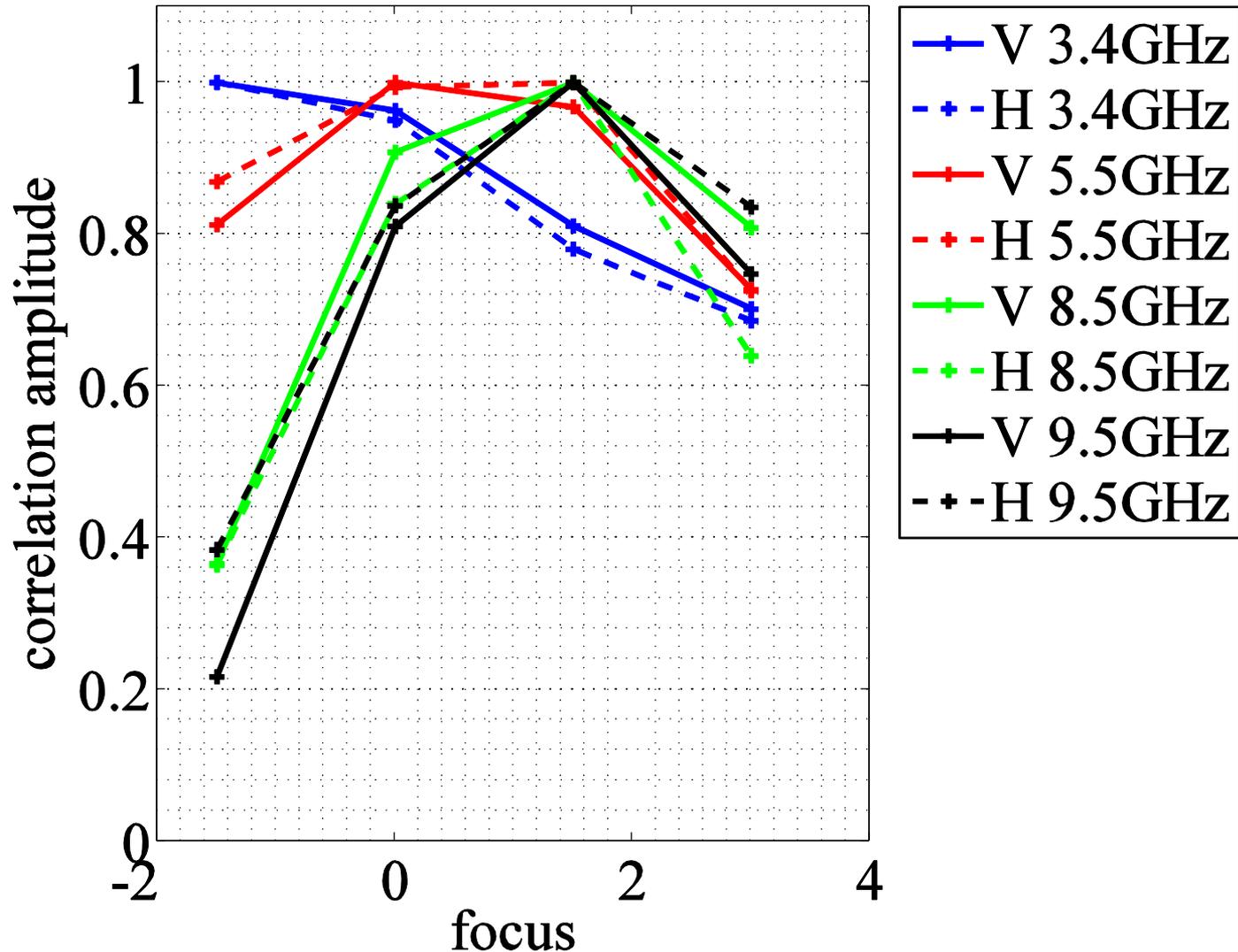
Proof of Concept System

- New equipment
 - Up Down Converter
 - New phase cal generator
 - Digital back end (DBE1 based on iBOB)
 - Mark5B+

Recent Results

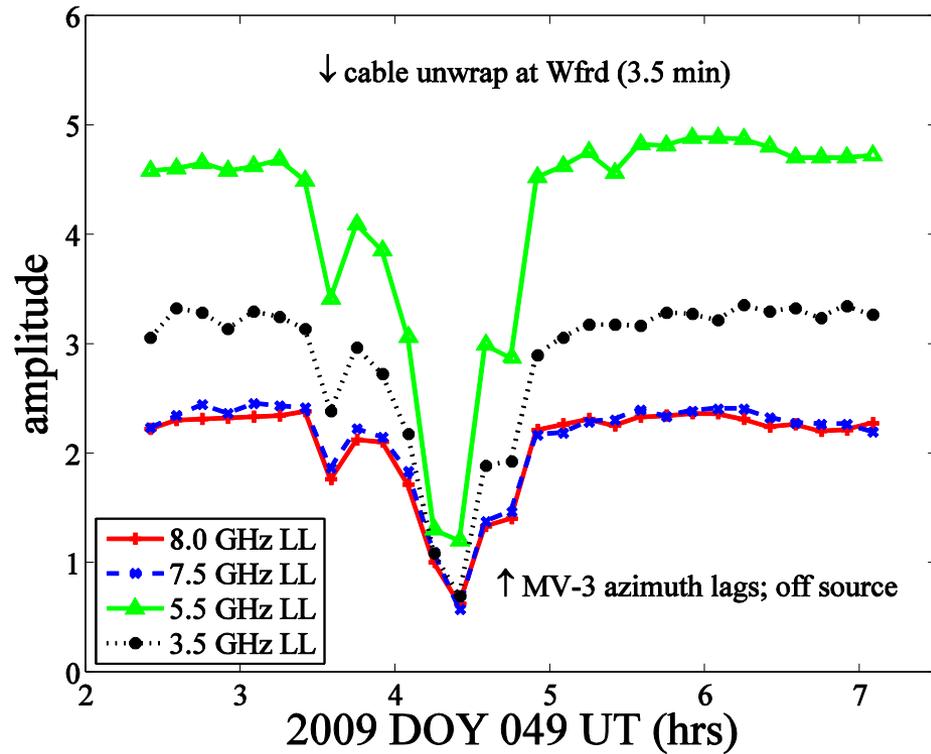
- Westford focus setting measurement
- Test sessions
 - Six hours with four bands identical frequencies
 - Five hours with bands contiguous -
 - Five hours on 4C39.25 polarization test

Westford focus measurement

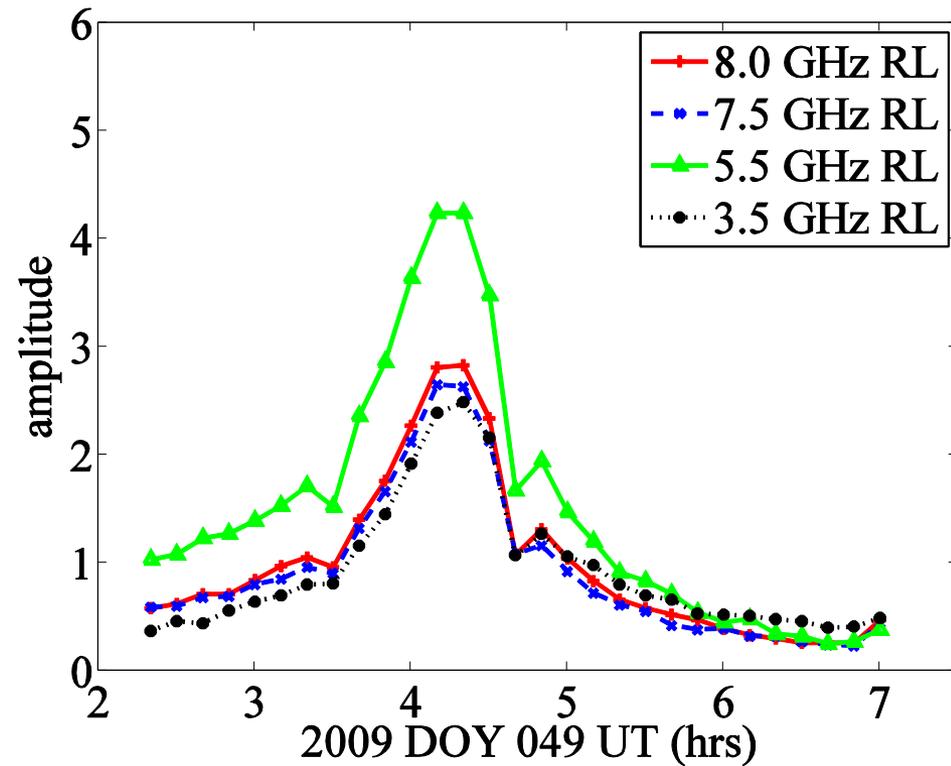


Polarization test

4C39.25 Wfrd-MV3



4C39.25 Wfrd-MV3



Proof of Concept System

- Problems/challenges encountered
 - RFI from TV at 500 MHz saturated LNAs
 - High-pass filter: > 3.1 GHz (cooled)
 - Protection diodes on LNA leads
 - Controlling four signal chains at two sites
 - Scripts developed for setting frequencies, levels
 - Phase cal processing in correlator
 - 5 MHz spacing and 32 MHz channel bandwidth
 - “Only software”; fixed

Proof of Concept System Status

	<u>Current</u>	<u>Next</u>
■ Feed	Lindgren	Eleven?
■ RF filter	>3.1 GHz	≥ 2.2 GHz
■ Phase cal	S/X pcal	“digital” pcal
■ DBE	DBE1(iBOB)	DBE2(ROACH)
■ Recorder	Mk5B+	Mk5C
■ Correlator	Mk4 (hardware)	Software (DiFX)



VLBI2010 Broadband System

On the antenna – cooled to $\sim 20\text{K}$

Feed

RF filters

Phase cal and noise diode injection

Low noise amplifiers

Send RF signals to the control room by optical fiber

In control room

Splitters to 4 bands (each polarization)

Flexible local oscillators

Digital back ends

Recorders or network connection

VLBI2010 Broadband System Characteristics

On the antenna – cooled to $\sim 20\text{K}$

Feed: 2 to 18 GHz

RF filters: >3.1 GHz

(falls off slowly above ~ 6 GHz)

Low noise amplifiers: 1 GHz to >12 GHz

In control room:

Flexible local oscillator (UpDown Converters):

1 to 13 GHz using 512 MHz Nyquist Zone filters

30 to 60 dB gain

Digital back end:

selects the eight odd 32 MHz channels from two 512 MHz inputs (V and H pol'n from one band)

Proof of Concept Demonstration

- Westford (Massachusetts):
 - 18m prime focus antenna
 - S/X receiver must be removed to install Dewar
- MV-3 (near Washington D.C.)
 - 5m Cassegrain antenna
 - Subreflector not matched to primary surface
- Signal chains are identical at both antennas
- All operations have been conducted from Haystack (except MV-3 antenna pointing)

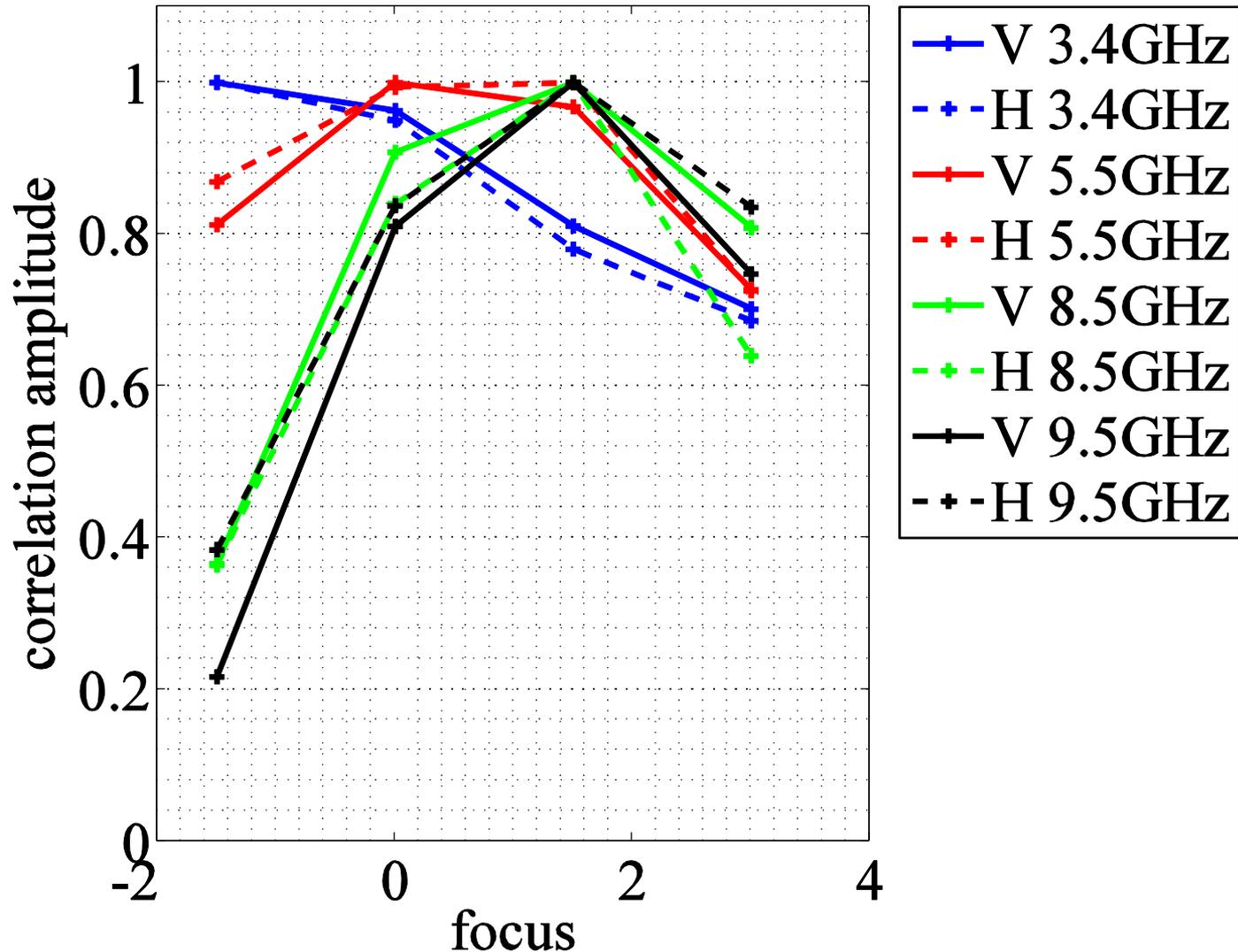
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Recent Results

- Westford focus setting measurement
- Test sessions
 - Six hours with four bands identical frequencies
 - Five hours with bands contiguous -
 - Five hours on 4C39.25 polarization test

Westford focus measurement



Polarization session

- Frequency coverage
 - 512 MHz bands
 - 3.5 GHz, 5.5 GHz, 7.5 GHz, and 8.0 GHz
- Source 4C39.25
 - Passes south of Westford and north of MV-3
 - Differential parallactic angle: +6 to +279
- Signal chains identical at both antennas
- All operations conducted from Haystack (except MV-3 antenna pointing)

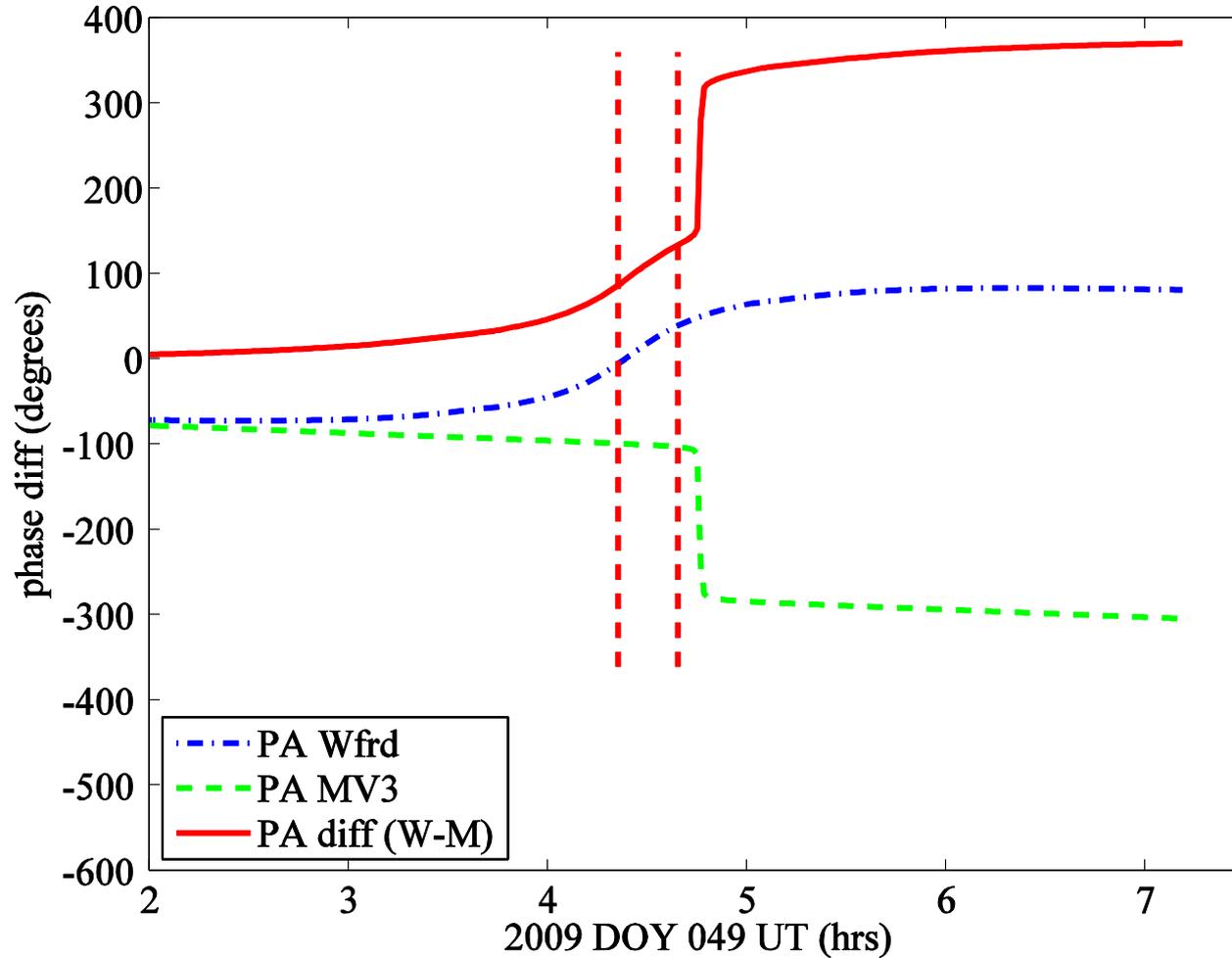
Results

- Expected SNR ~ 250
 - BW = 256 MHz
 - Integration time = 570s
 - Correlated flux density = 10Jy
 - SEFDs = 3000Jy/100,000JY
 - Observed SNR (max)

2009-040 freq (GHz)	3.5	5.5	7.5	8.0
	140	200	120	120

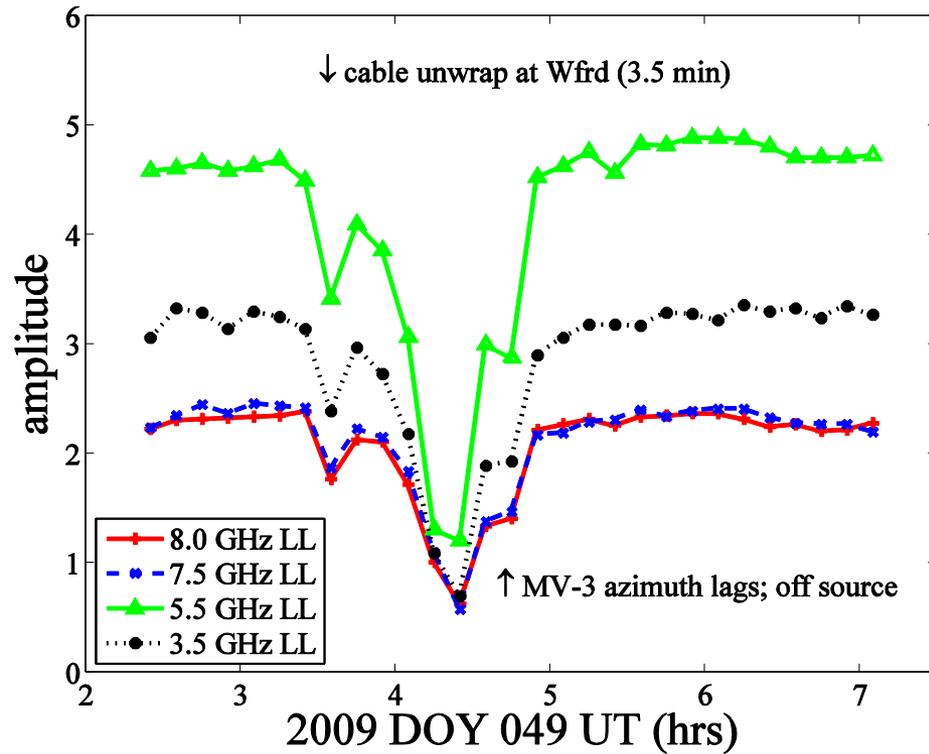
Parallactic Angle Differences

4C39.25 Wfrd-MV3

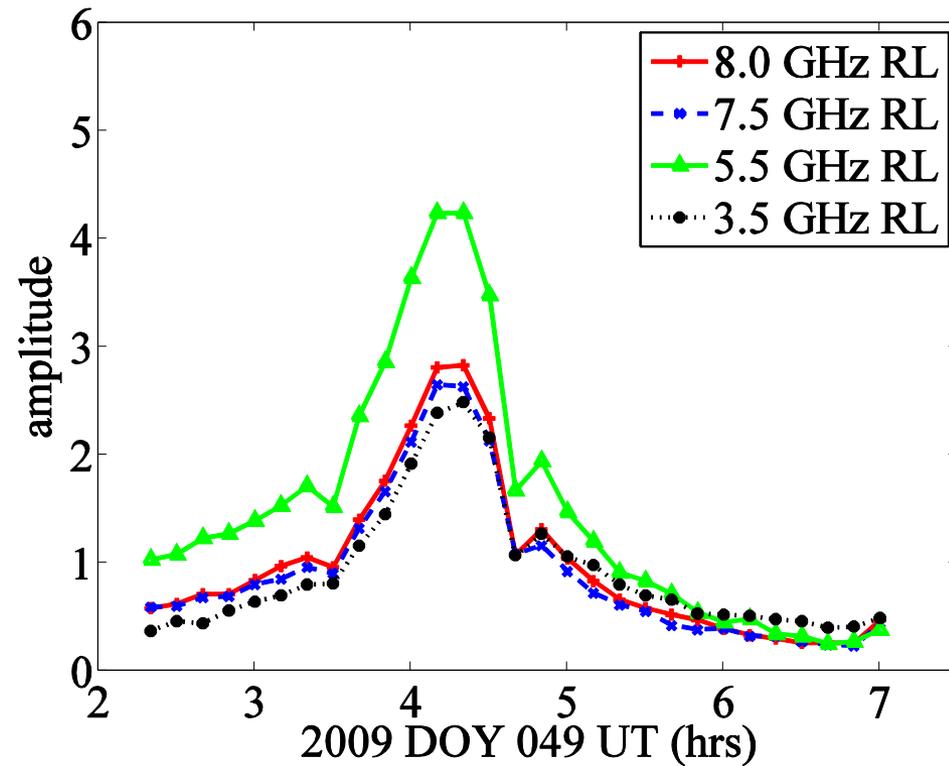


Polarization test

4C39.25 Wfrd-MV3

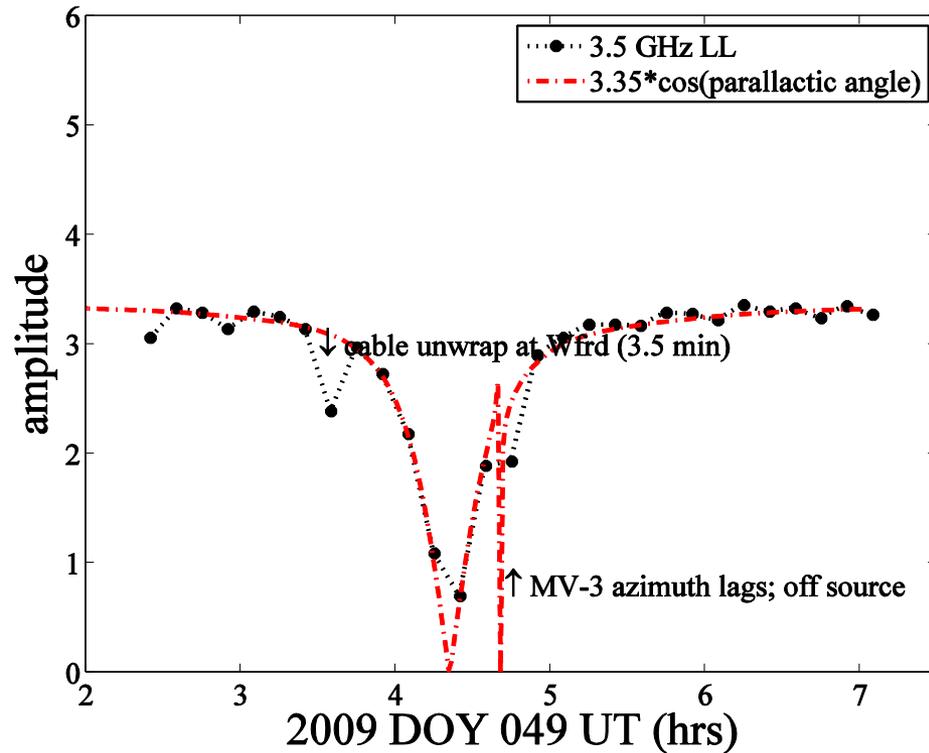


4C39.25 Wfrd-MV3

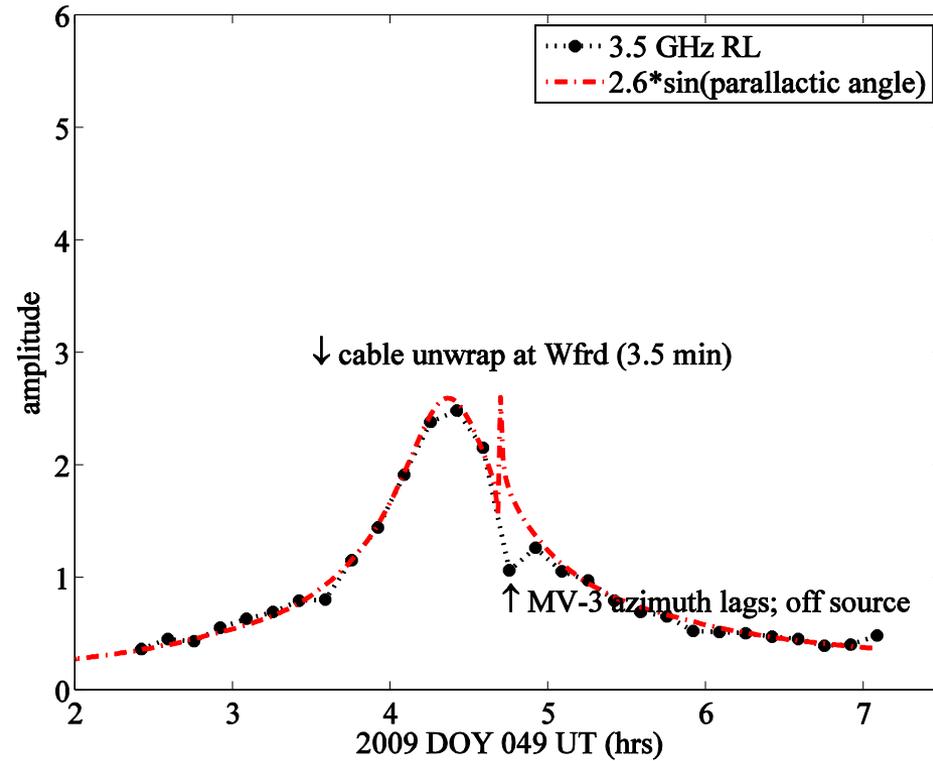


Polarization test

4C39.25 Wfrd-MV3



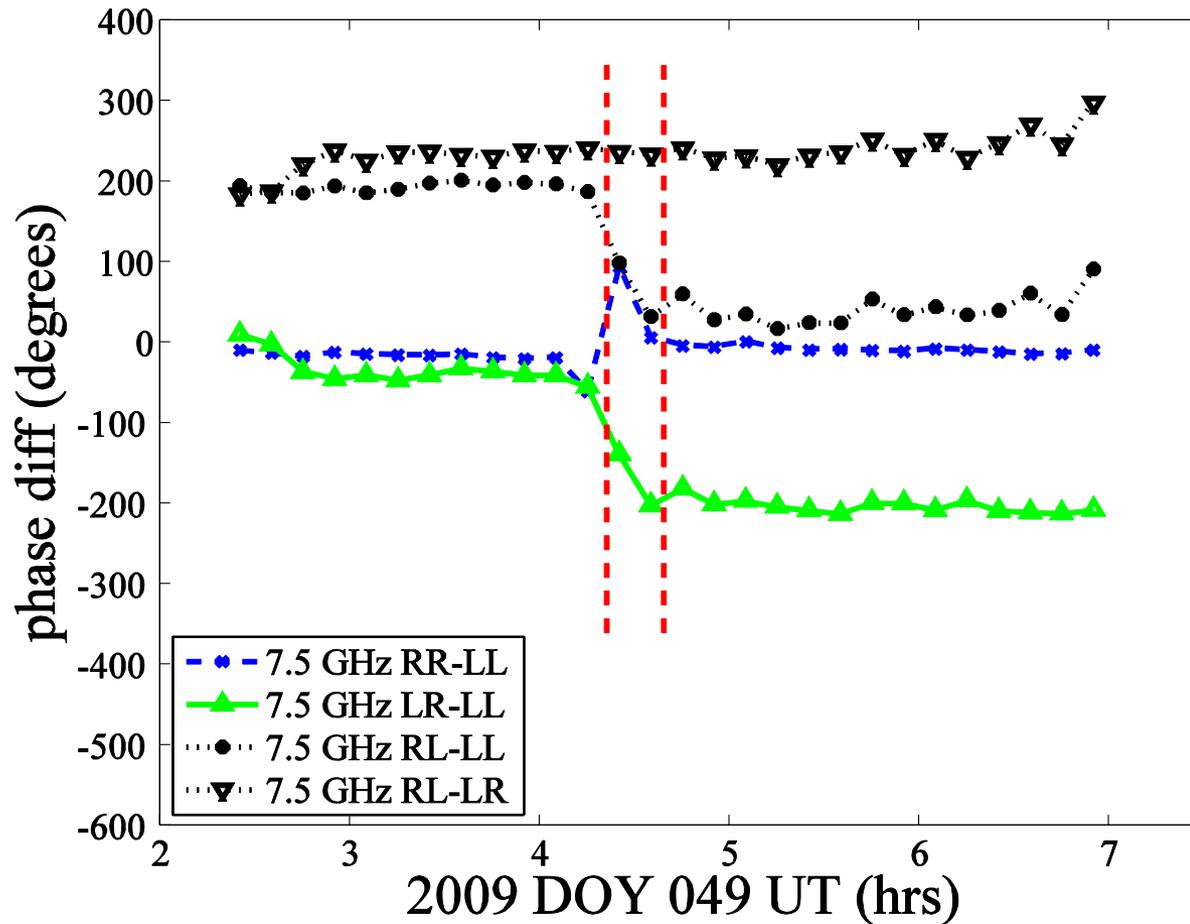
4C39.25 Wfrd-MV3



Polarization Phase Differences

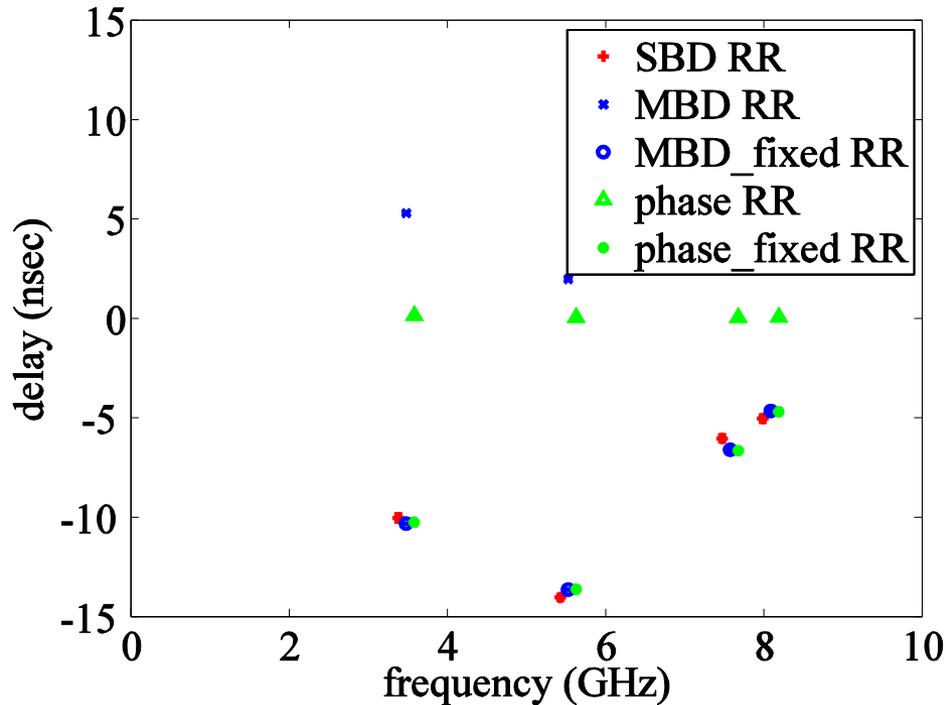
(no phase cal applied)

4C39.25 Wfrd-MV3

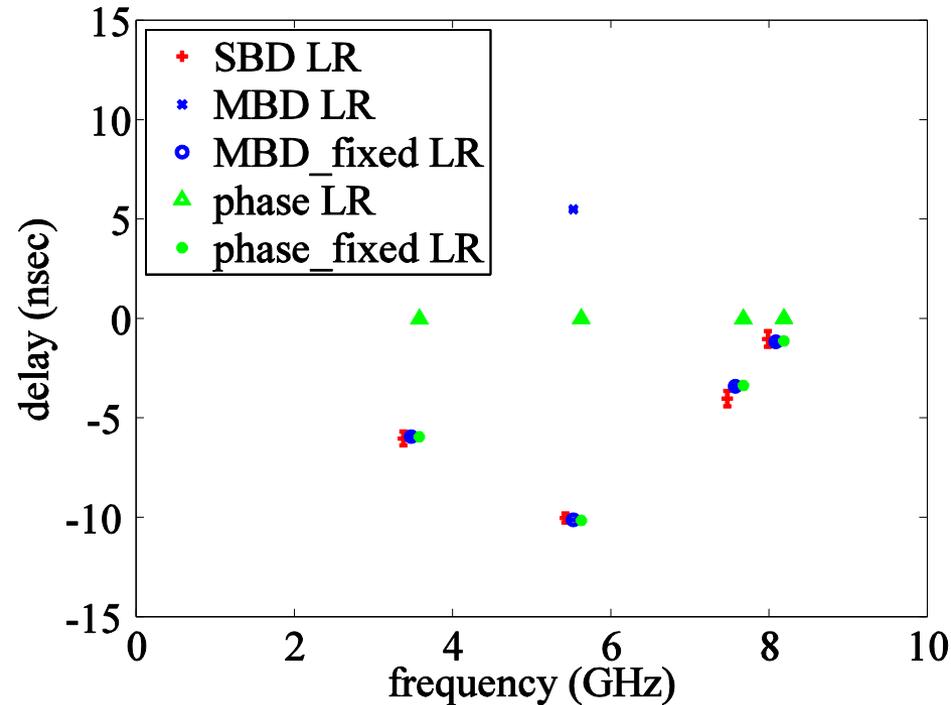


Polarization test (no phase cal applied)

4C39.25 Wfrd-MV3

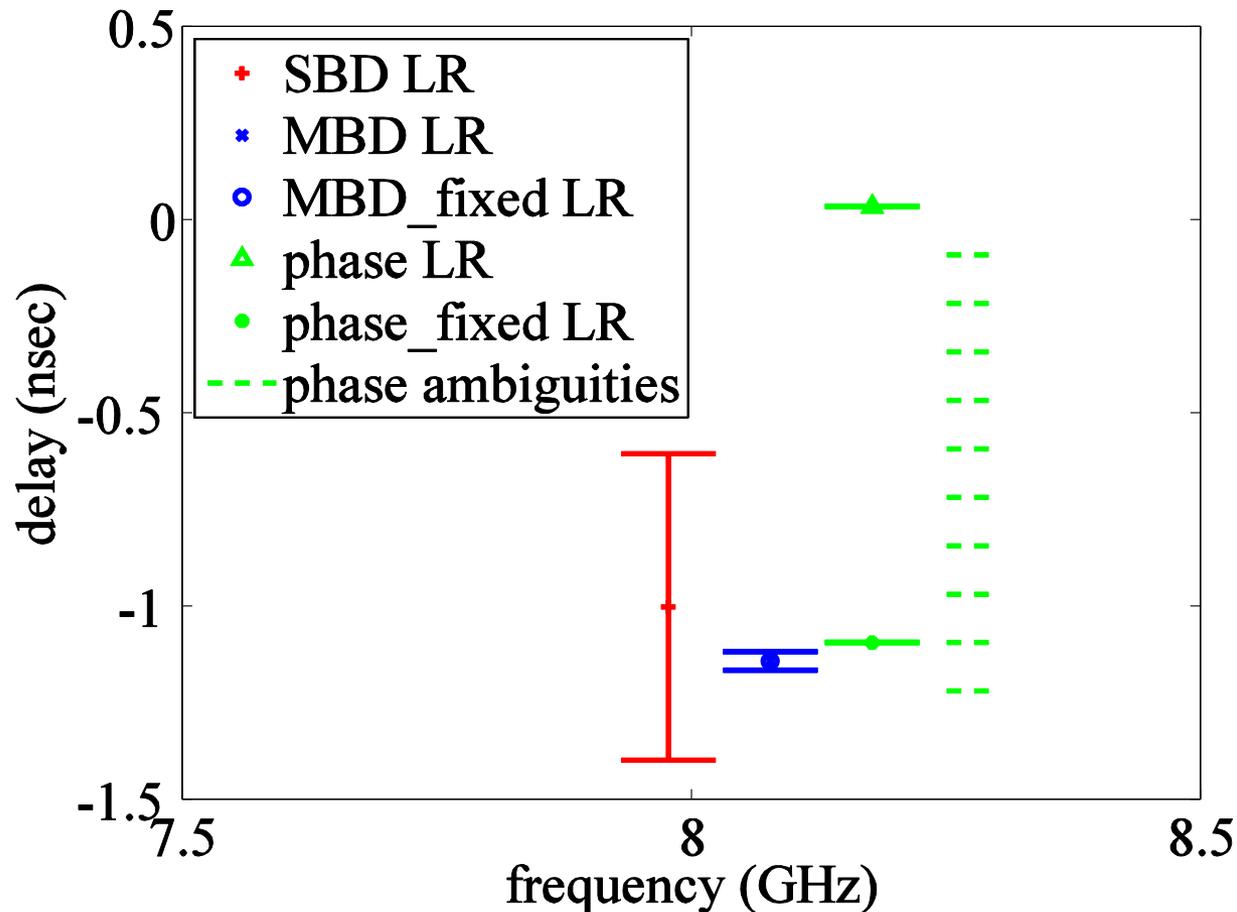


4C39.25 Wfrd-MV3



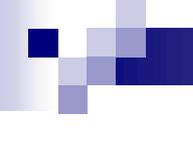
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Benefits of Ku-Ka Band Reception

- Advantages

- Sources more compact
- Possibility of observing as a co-pointing WVR

Ku to Ka Band Reception

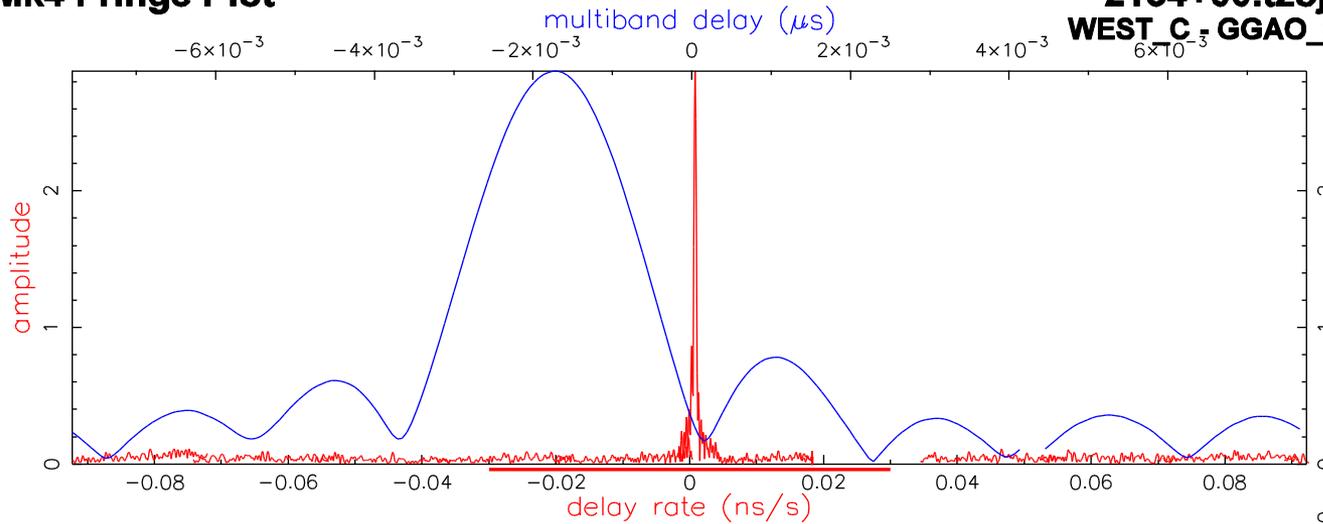
■ Disadvantages

- Antenna tolerances more exacting
 - Surface accuracy
 - Pointing
- Antenna sensitivity lower
 - T_{sys} higher
 - Atmosphere contribution higher
- Fewer sources with same correlated flux density on a given baseline

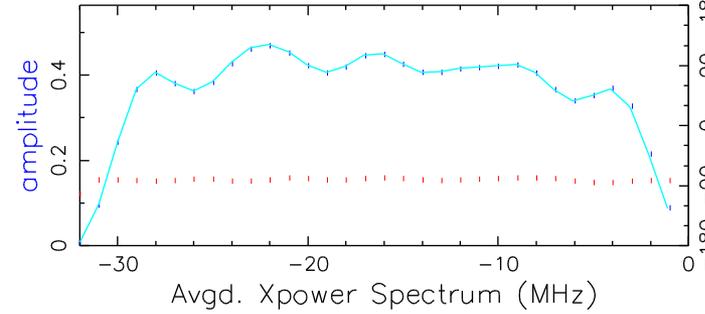
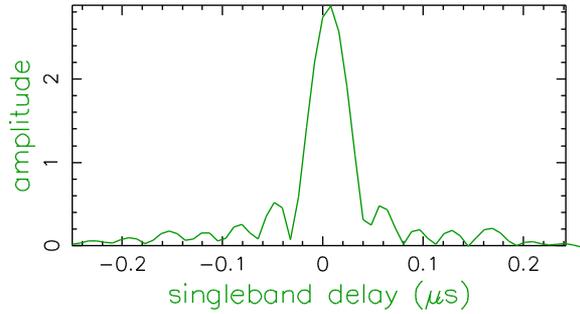


Mk4 Fringe Plot

2134+00.tzobjj, 034-2010, CI
 WEST_C_3 GGAO_C, fgroup X, pol LL

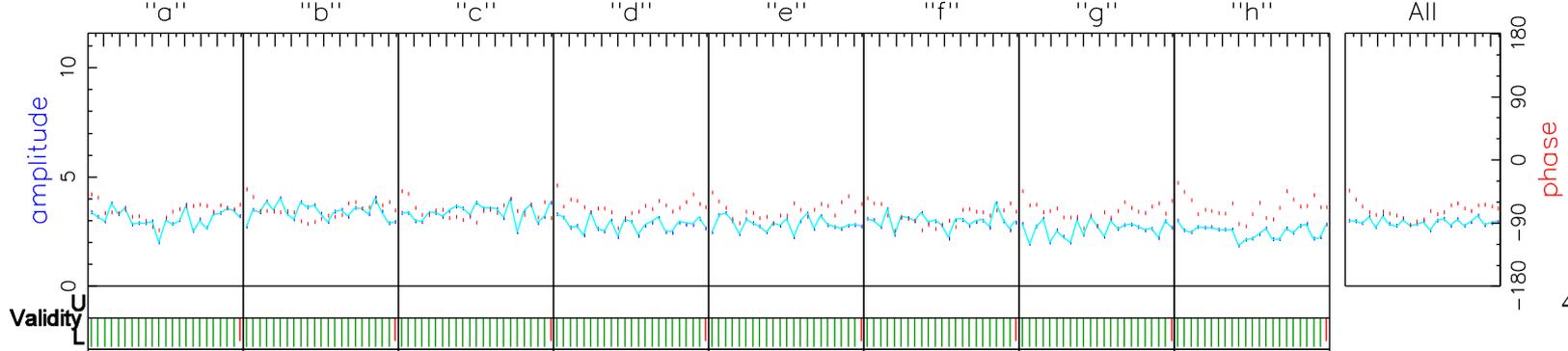


Fringe quality 9
SNR 122.9
PFD 0.0e+00
Intg.time 569.690
Amp 2.890
Phase -71.8
Sbdelay (us) -0.001567
Mbdelay (us) -0.001567
Fr. rate (Hz) 0.003109
Ref freq (MHz) 5416.4000
AP (sec) 1.000



Exp. X4 tst
Exper # 3255
Yr:day 2009:034
Start 201001.00
Stop 201931.00
FRT 201516.00
Corr. date: 2009:035:001550
Fourfit date: 2009:035:003653
Position (J2000) 21h36m38.5863s +0°41'54.213"

Amp. and Phase vs. time for each freq., 23 segs, 25 APs / seg (25.00 sec / seg.), time ticks 30 sec



VLBI2010 Broadband System

- Expected SNR for observed SEFDs (256 MHz, 570s, 10Jy, 3KJy, 100K)

$$\text{SNR} = 0.8 * \sqrt{2 * B * T} * S_c / \sqrt{\text{SEFD_Wf} * \text{SEFD_Gg}} = 250$$

2009-DOY040 freq	3.5	5.5	7.5	8.0
Observed SNR (max)	140	200	120	120

1. New equipment built especially for VLBI2010/uvlbi
 1. UDCs
 2. DBEs – modified for dual polarization recording
 3. Digital PCAL generator

Westford focus setting measurement

Test sessions

Dewar (1)

- Feed (for initial operations of VLBI2010)
 - 2.2 to ~14 GHz (possibly to 36 GHz eventually)
 - Phase center and beamwidth should be frequency independent (implies dual linear polarization now)
- Phase cal and noise cal
 - Signals split and injected to both polarizations following feed but before LNAs
 - Possibly external phase cal signal also for feed cal

Dewar (2)

- RF filter (if necessary)
 - TV at 500 MHz saturates unprotected LNAs
 - Cell phones about 2 GHz may be problem
 - Communication satellites above 10 GHz (but positions known)
 - Other ?
- Low Noise Amplifiers (one for each pol'n)
 - 2.2 to ~14 GHz

Backend (control room) - 1

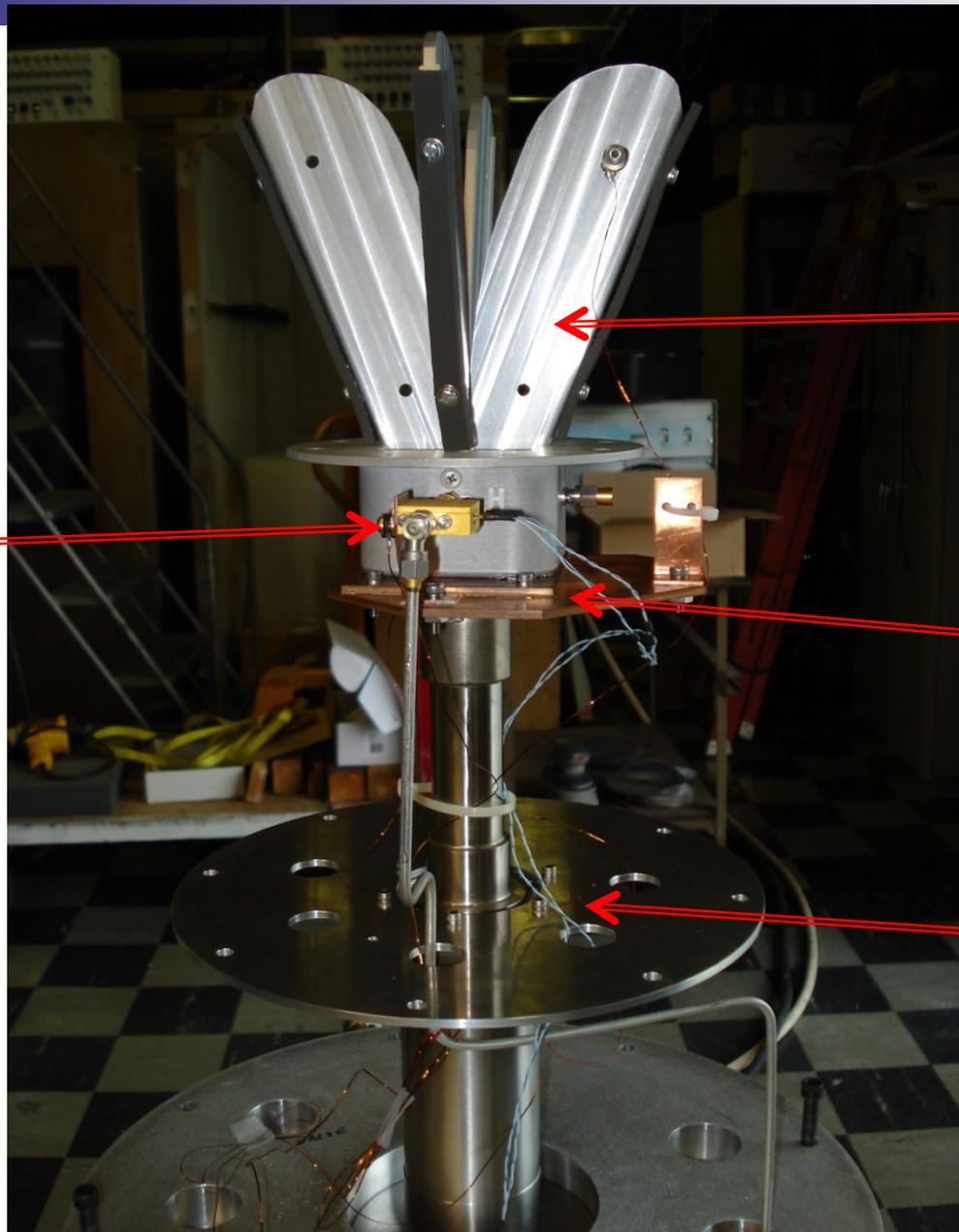
- RF carried to control room via optical fiber
- Optical fiber receivers, amplifiers, 4-way splitters (both polarizations) in one chassis
- RF is down-converted directly to video
 - Alan Rogers' design also amplifies and NZ filters

Backend (control room) - 2

- Digital Back End
 - Analog-to-digital conversion
 - Poly phase filter and freq channel selection
or
 - Baseband converters (DBBC)
- Data storage
 - Directly to disks
or
 - Network to intermediate storage or correlator
- Real-time e-vlbi



LNA



Lindgren feed
(linear pol'n)

Cryo
Refrigerator
20K Station

77K Station

12 m Antenna

